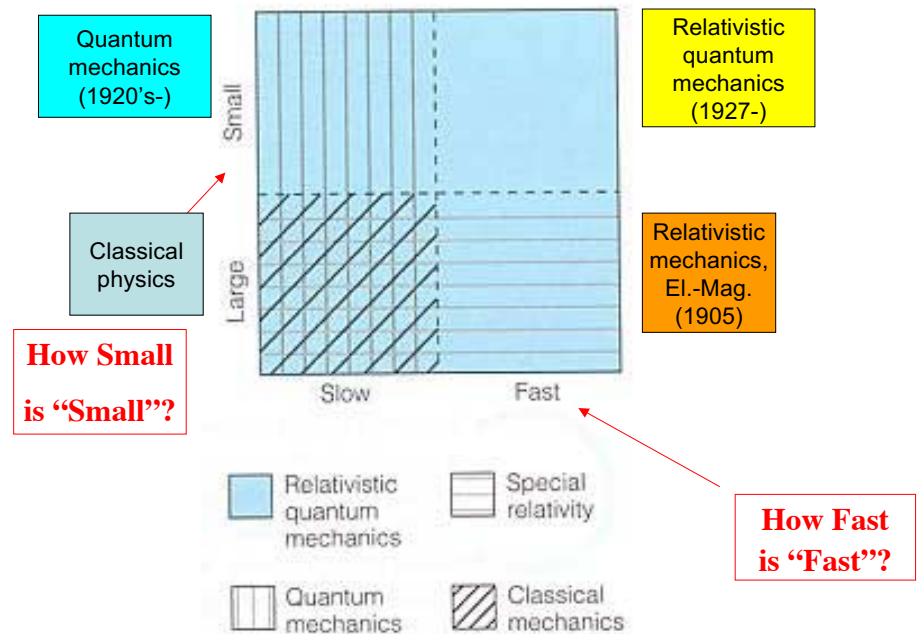
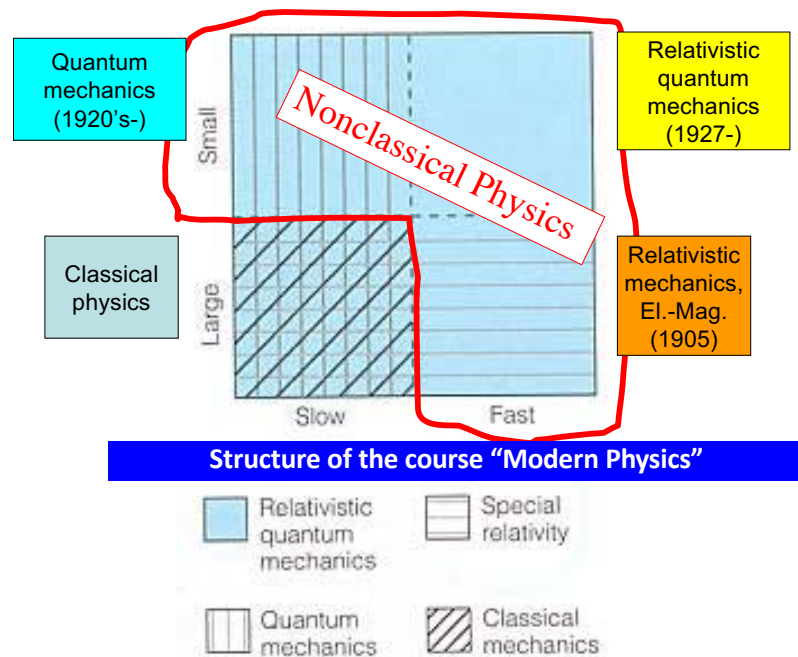
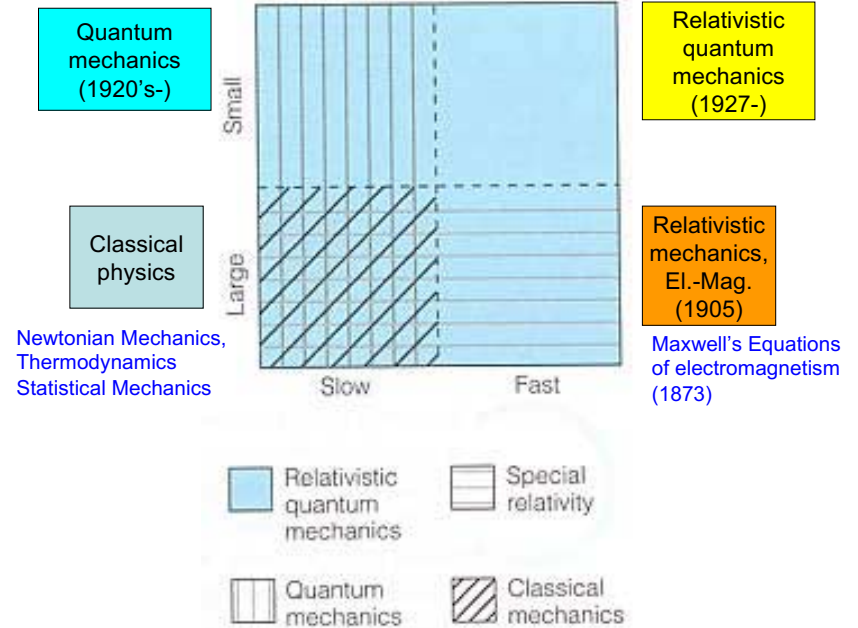


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Lecture 2

Aug. 27, 2024



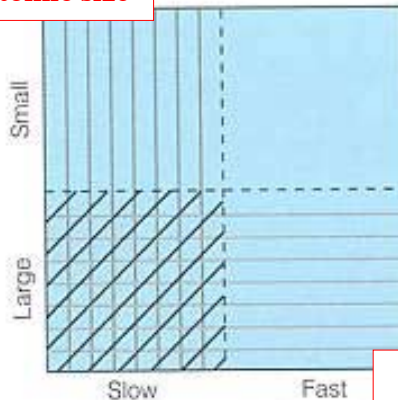
Small => e.g. atomic size

Quantum mechanics
(1920's-)

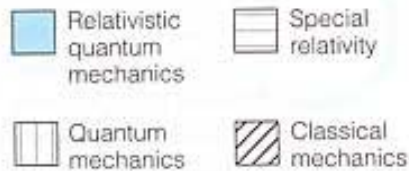
Relativistic quantum mechanics
(1927-)

Classical physics

Relativistic mechanics, El.-Mag.
(1905)

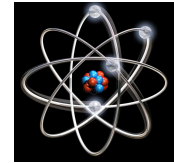
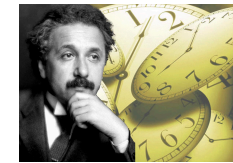


Fast => $v \sim c$
 $c = \text{the velocity of light}$



Modern Physics

- Two basic ideas
 - Time and space are not absolutes.
 - **Particles behave like waves and waves behave like particles.**
- Two branches
 - Special Relativity
 - Quantum Mechanics



With an understanding of these branches, we can then explore areas of modern physics such as **superconductivity**, **modern optics**, **nuclear physics**, **particle physics** and **cosmology** - along with a host of other areas of science.

6

Chapter. 3 Wave & Particles I

EM-“Waves” behaving like “Particles”

Outline:

- Blackbody Radiation (Plank; 1900; 1918*)
- The Photoelectric Effect (Einstein; 1905; 1921*)
- The Production of X-Rays (Rontgen; 1901; 1901*)
- The Compton Effect (Compton; 1927; 1927*)
- Pair Production (Anderson; 1932; 1936*)
- Is It a Wave or a Particle? → Duality?

Chapter. 4 Wave & Particles II

“Matter” behaving as “Waves”

Outline:

- A Double-Slit Experiment (watch “video”)
- Properties of Matter Waves
- The Free-Particle Schrödinger Equation
- **Uncertainty Principle**
- **The Bohr Model of the Atom**
- **Mathematical Basis of the Uncertainty Principle – The Fourier Transform**

Chapter. 5

Bound States: Simple Case

Purpose:

- To make QM useful in real application,
- we must have a way to account for the effects of external forces**

Let's start with the Schrödinger eq. to include these effects.

** interaction of object with its surrounding

Chapter. 5

Bound States: Simple Case

Outline:

- The Schrödinger Equation (for interacting particles)
- Stationary States
- Physics Conditions: Well-Behaved Functions
- A Review of Classical Bound States
- Case 1: Particles in a Box – The Infinite Well
- Case 2: The Finite Well
- Case 3: The Simple Harmonic Oscillator
- Expectation Values, Uncertainties, and Operators

Chapter. 6

Unbound States

Outline:

- The Potential Step
- The Potential Barrier & Tunneling
- Alpha Decay & Other Applications
- Particle-Wave Propagation

Chapter. 7

QM in 3-dims & Hydrogen Atom

Outline:

- The Schrödinger Eq. in 3-Dimensions
- The 3D Infinite Well
- Energy Quantization & Spectral Lines in Hydrogen
- The Schrödinger Eq. for a Central Force
- Angular Behavior in a Central Force
- The Hydrogen Atom
- Radial Probability
- Hydrogen-like Atoms

Chapter. 8

Spin & Atomic Physics

Outline:

- Evidence of Angular Momentum Quantization
- Identical Particles
- The Exclusion Principle
- Multi-electron Atoms & the Periodic Table
- Characteristic X-Rays

It's open said that in Q.M. there're only 3 bound-state problems solvable (w/o numerical approximation tech.)

1. Infinite well, 2. Harmonic oscillation, 3. hydrogen atom
- all 1-particle problem.

Most real application: multiple system. so, let's start an atom with multiple electrons

Chapter. 9

Statistical Mechanics

Statistical mechanics is NOT non-classical or modern physics in the same sense that special relativity and QM are. Rather, it's a distinct area of physics that applies to many others; either classical or QM.

Q.: Why study it now? A.: many modern physics require it.

e.g. A gas laser: thermodynamics system of gas molecules

A semiconductor: thermodynamics system of atom bound in a solid lattice

**** Thermodynamic system = countless particles; precise average behavior.**

Chapter. 9

Statistical Mechanics

Outline:

- Historical Overview
- The Boltzmann Distribution
- Maxwell Velocity Distribution
- Equipartition Theorem
- Maxwell Speed Distribution
- Classical and Quantum Statistics
- Fermi-Dirac Statistics
- Bose-Einstein Statistics

Chapter. 10

Molecules and Solids

Outline:

- 10.1 Molecular Bonding and Spectra
- 10.2 Stimulated Emission and Lasers
- 10.3 Structural Properties of Solids
- 10.4 Thermal and Magnetic Properties of Solids
- 10.5 Superconductivity
- 10.6 Applications of Superconductivity

In Chapter 7 & 8, we learned about the properties of individual atoms. This chapter builds on that knowledge to find out what happens when those atoms join together to form molecules & solids.

Chapter. 11

Semiconductor Theory and Devices

Outline:

- 11.1 Band Theory of Solids
- 11.2 Semiconductor Theory
- 11.3 Semiconductor Devices
- 11.4 Nanotechnology

In Chapter 10 you learned about structural, thermal, and magnetic properties of solids. In this chapter we concentrate on electrical conduction.

Historical Development

Newton(1704): light as a stream of particles.

Descartes (1637), Huygens, Young, Fresnel (1821), Maxwell: by mid-19th century, the wave nature of light was established (*interference* and *diffraction*, transverse nature of EM-waves).

Physics of the 19th century: mostly investigation of *light waves*

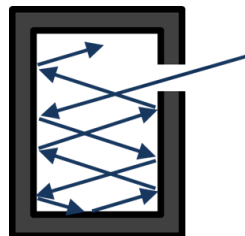
Physics of the 20th century: *interaction of light with matter*

One of the challenges – understanding the “*black body spectrum*” of thermal radiation

Black body:

In physics, a black body is an idealized object that absorbs all incident E&M radiation. No E&M radiation passes through the black body and none is reflected.

Because no light is reflected or transmitted, the object appears black when it is cold.



An approximate realization of a black body as a tiny hole in an insulated enclosure

Chapter. 3

Wave & Particles I

EM-“Waves” behaving like “Particles”

Outline:

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Black body:

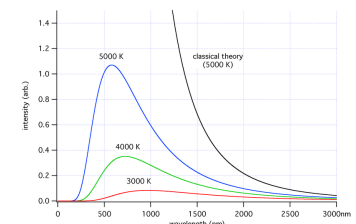
In physics, a black body is an idealized object that absorbs all incident E&M radiation. No E&M radiation passes through it and none is reflected.

Because no light is reflected or transmitted, the object appears black when it is cold.

However, a black body emits a temperature-dependent spectrum of light. (see Fig.)

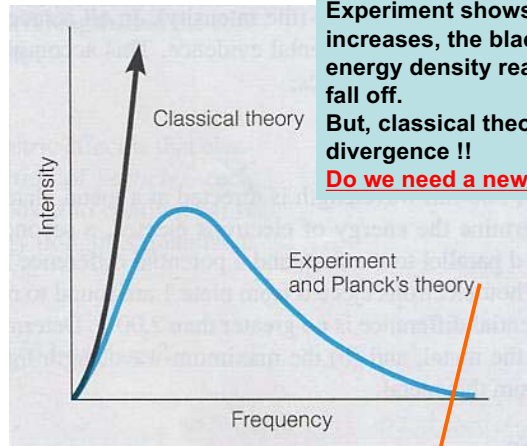
This thermal radiation from a black body is termed black-body radiation.

As the temperature decreases, the peak of the black-body radiation curve moves to lower intensities and longer wavelengths.



Black Body Radiation

(Max Planck 1900)



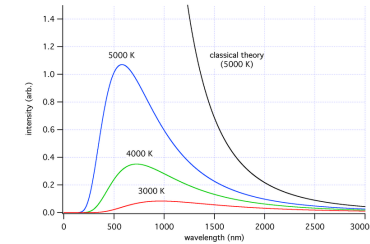
Experiment shows that as frequency increases, the blackbody spectral energy density reaches a max. then fall off.

But, classical theory predicts a divergence !!

Do we need a new theory?

(More in Appendix C)

Historical Development



In 1900, **Planck** suggested a solution based a revolutionary new idea:

Emission and absorption of E&M radiation by matter has **quantum nature**:

i.e. the energy of a quantum of E&M radiation emitted or absorbed by a harmonic oscillator with the frequency f is given by the famous Planck's formula

$$E = hf$$

, where h is the Planck's constant

$$h \approx 6.626 \times 10^{-34} \text{ J} \cdot \text{s}$$

- at odds with the "classical" tradition, where energy was always associated with amplitude, not frequency

Also, in terms of the angular frequency

$$\omega = 2\pi f$$

$$E = \hbar \omega$$

where $\hbar = \frac{h}{2\pi}$

$$\hbar \approx 1.05 \times 10^{-34} \text{ J} \cdot \text{s}$$

The Planck's Black-Body Radiation Law:

The Energy (E) in the electromagnetic radiation at a given frequency (f) may take on values restricted to

$$E = nhf$$

where:

n = an integer

h = a constant $h \approx 6.626 \times 10^{-34} \text{ J} \cdot \text{s}$

("Planck Constant")